## Development of Integrated Receiver for Radio Astronomy

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The 10th Workshop on Submm Wave Receiver Technologies in Eastern Asia

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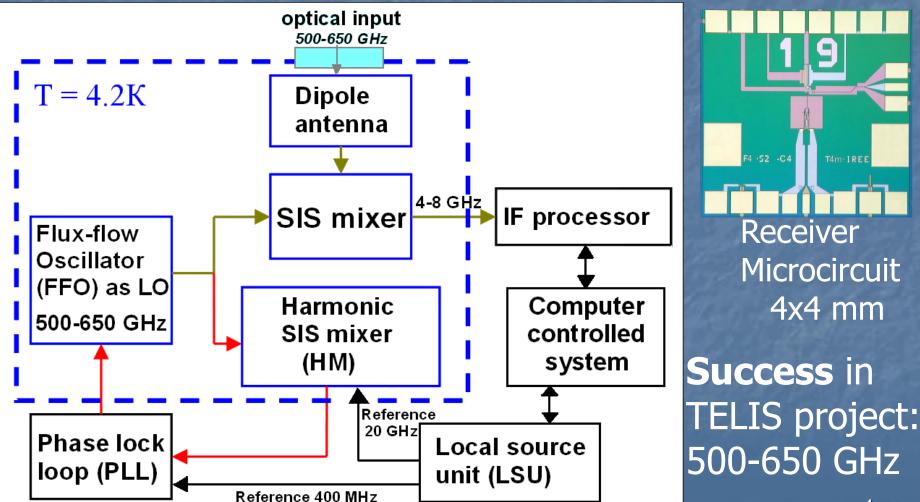
Wuxi, China

## Development of Integrated Receiver for Radio Astronomy

- Superconducting Integrated Receiver of 500-650 GHz
- New applications of Integrated Receiver
   Development of flux-flow oscillator (FFO) for 250-400 GHz
- Development and testing circuit including FFO, SIS and Harmonic Mixers.
   Calculation and experimental results
   Summary

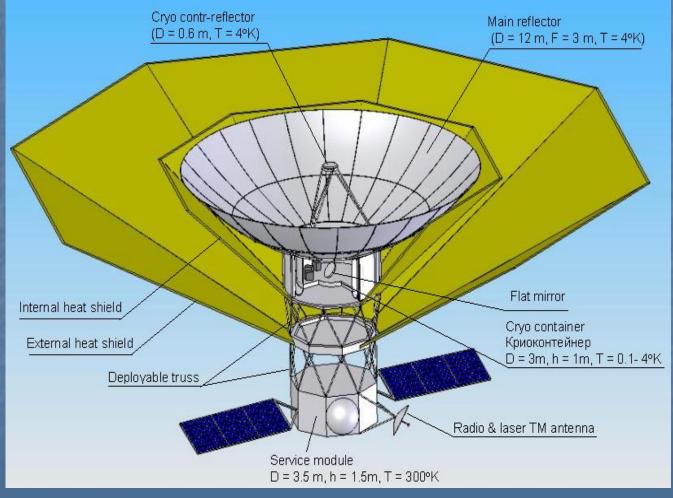
## Superconducting Integrated Receiver

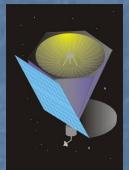
## Superconducting Integrated Receiver with phase-locked FFO



## New applications of Integrated Receiver

### New applications of Integrated Receiver (1) **"Millimetron"** – Russian Space Agency ( > 2017) 12 m cryogenic mirror; $\lambda = 0,01$ - 20 mm.





#### Ground-space interferometer



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#### New applications of Integrated Receiver (2)

ESPRIT – Exploratory Submm Space Radio-Interferometric Telescope



## The six elements of ESPRIT in an Ariane 5

Telescope sizes ~ 3.5 meter ; off-axis Number of elements N = 6 (15) baselines) Projected baselines 200 - 1000 meter Frequencies: Spots in the range 0.5 – 6 THz Front Ends - (0.5 – 1.5 THz): SIS mixers, multiplier LO / SIR = FFO + SIS + HM (1.5 – 6 THz) HEB mixers, QCL as LO System temperature < 1000 KIF bandwidth > 4 GHz (goal 8 GHz) 

#### New applications of Integrated Receiver (3)

Submillimeter Telescope (SMT) Arizona Radio Observatory (ARO)



Main reflector: paraboloid D=10 m; F/D=0.35. Subreflector: d=0.69 m; **SIS-490:** The SORAL 490 GHz single-channel receiver; T(DSB) = 110-150 K across its 425 to 500 GHz tuning range.

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#### New applications of Integrated Receiver (4)

30-cm **POrtable Submillimeter Telescope** (POST) *Purple Mountain Observatory (PMO); Nanjing.* Site: Delingha of Qinghai province *(altitude ~3200 m)* 



#### **Required parameters:**

Frequency - 345 GHz Tr (DSB) ≤ 100 K

IF range - 3.6-4.6 GHz Spectral resolution  $\leq$  30 KHz

Output power  $\geq$  5 dBm Output frequency - 0.01-1GHz

Dissipated power at 4.2K stage ≤ 100 mW

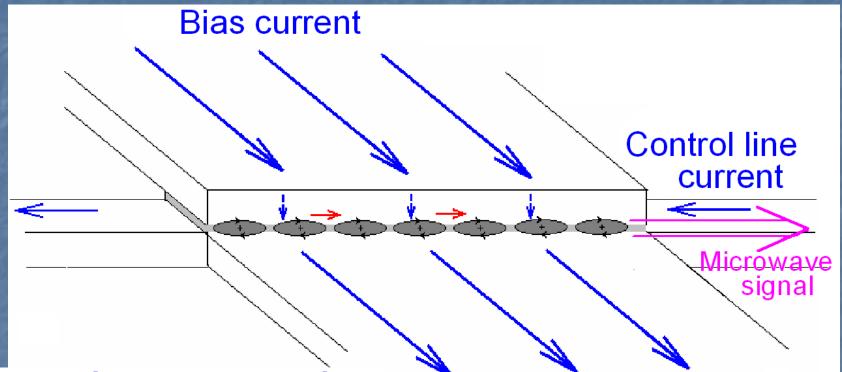
## **Integrated Receiver of 345 GHz for POST:**

collaboration:
 IREE, Moscow, Russia
 PMO, Nanjing

**Development of such Integrated Receiver for POST includes stages:** New FFO design for 325-365 GHz Microcircuit design to pump SIS and Harmonic mixer with FFO power Integrated lens-antenna Mechanical design for mounting a new receiver at the telescope Related stuff and equipment (cables, amplifiers, controlling software, etc.)

## **Development of Flux-flow oscillator for 250-400 GHz**

## Flux-flow oscillator as LO



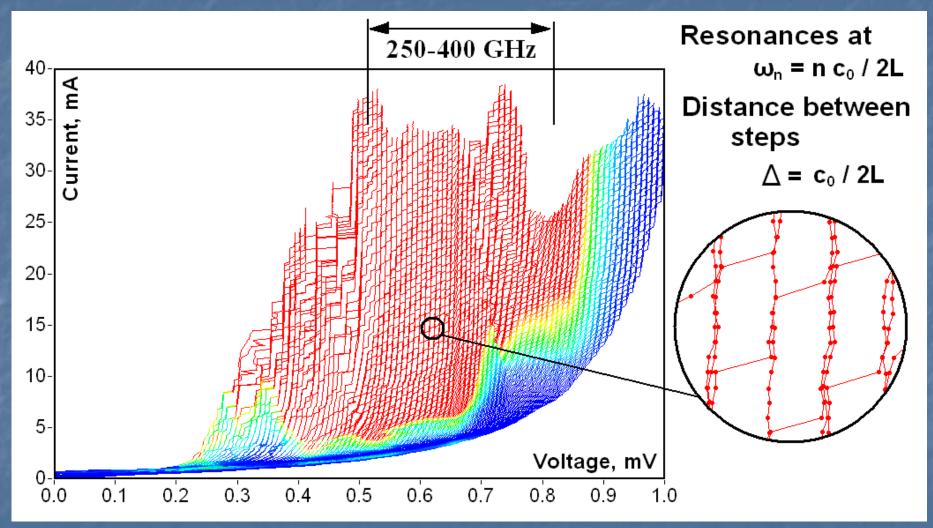
#### Josephson equation

$$f_{\rm FFO} = \frac{2eV_{\rm FFO}}{hc}$$
(483.6 GHz/mV)

Working regimes:

 400-650 GHz – Fiske steps and flux-flow regime
 250-400 GHz – Fiske steps regime with low damping

## IV-curves of FFO



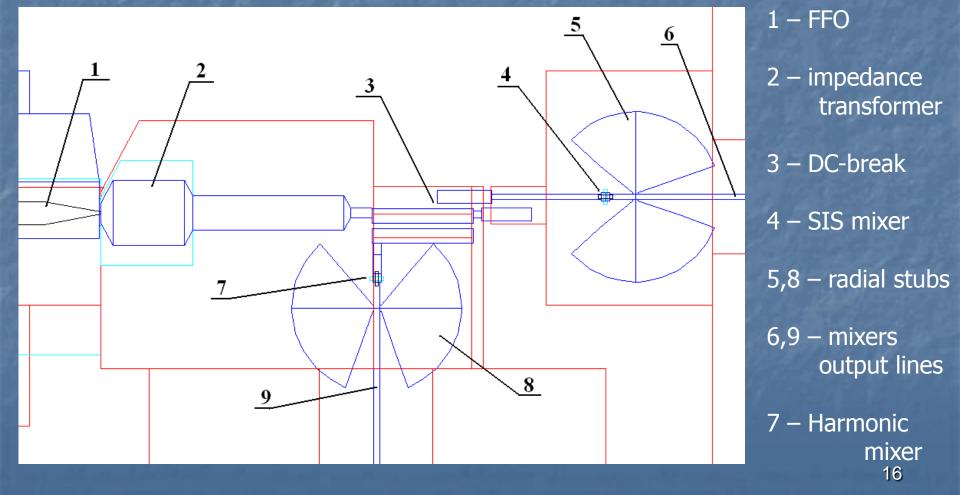
FFO Length 700 µm was taken

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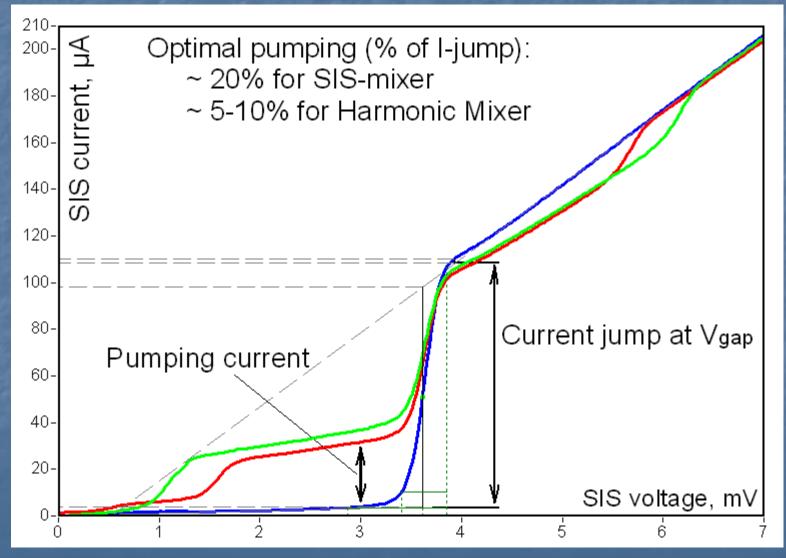
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## Development and testing of integrated circuit

# Development of testing microcircuit FFO + SIS mixer + Harmonic mixer Power matching of FFO with mixers is required

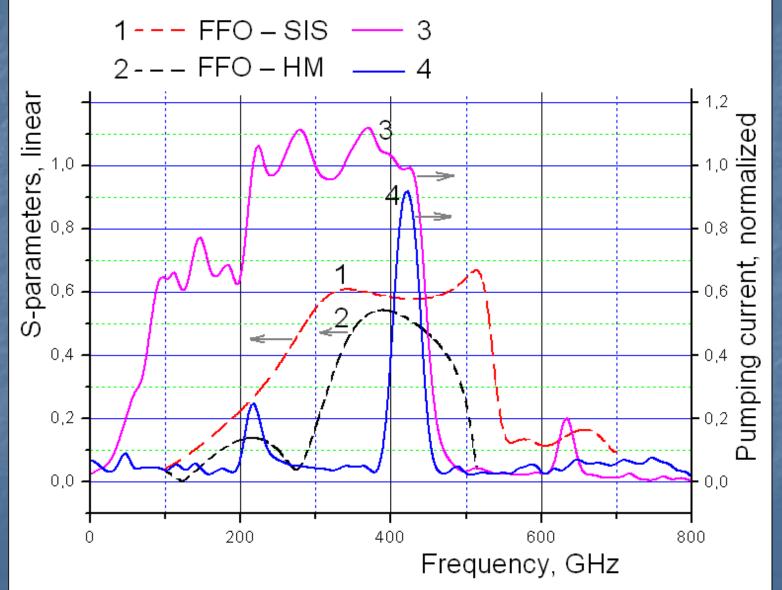


## <u>IV-curves of SIS mixer</u> showing pumping by FFO power



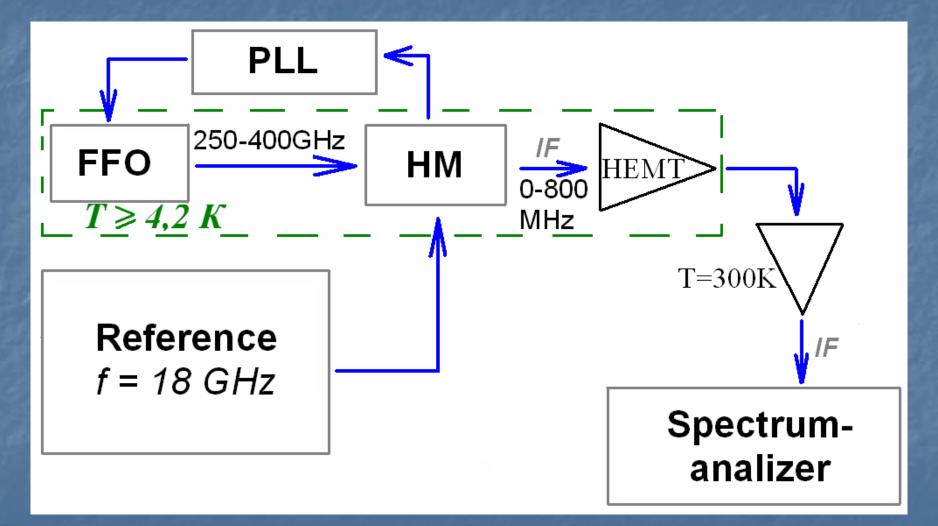
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## FFO and mixers power coupling



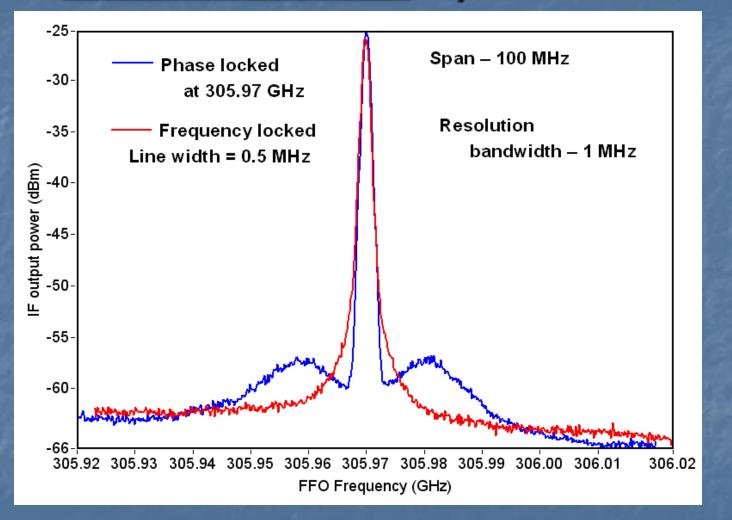
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## FFO Linewidth measurements



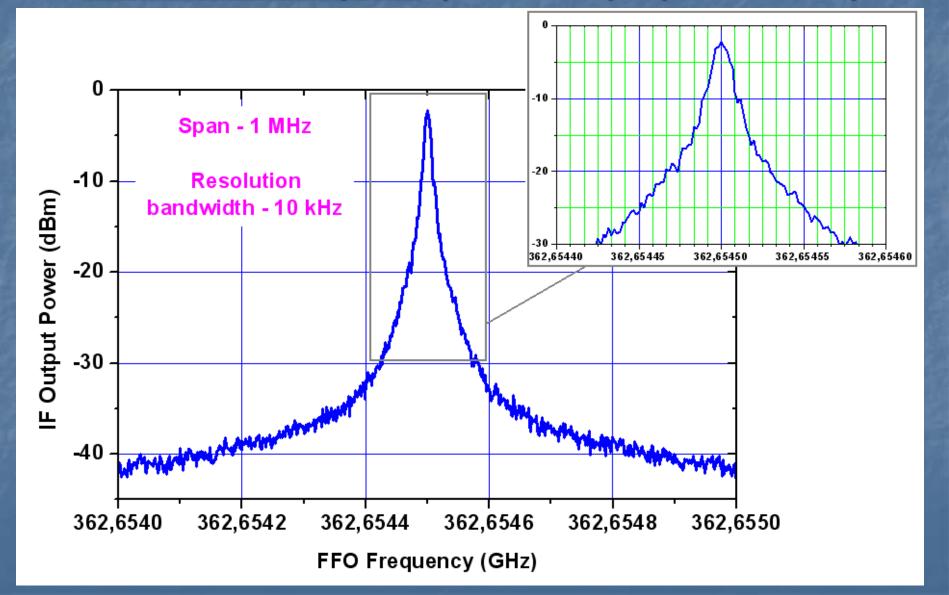
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## FFO radiation spectrum



#### Linewidth at best points below 1 MHz in the range 325-365 GHz 15-18 November, 2009 Wuxi, China Was obtained. 20

#### <u>Measured</u> by integrated receiver a <u>signal of</u> <u>external multiplier</u> (driven by synthesizer)



## Summary

FFO linewidth below 1 MHz was demonstrated in range 325-365 GHz Sufficient power matching between FFO and both mixers (SIS and HM) was obtained What has to be done: Integrated lens-antenna to be developed Circuit design with antenna to be retreated FFO design to be improved Mechanical construction for mounting receiver into telescope to be realized

## Thanx for your attention :)

## Phase Noise of the PL FFO

