## Optical design and verification of the sub-millimeter limb sounder TELIS

P. Yagoubov<sup>1</sup>, G. de Lange<sup>1</sup>, A. Baryshev<sup>2</sup>, R. Hesper<sup>2</sup>, V. Koshelets<sup>3</sup>, G. Wagner<sup>4</sup>, M. Birk<sup>4</sup>, A. Murk<sup>5</sup> TELIS



<sup>1</sup>National Institute for Space Research, SRON, the Netherlands <sup>2</sup>Kapteyn Astronomical Institute /SRON <sup>3</sup>Institute of Radio Engineering and Electronics, IREE, Russia <sup>4</sup>Institute for Remote Sensing Technology, DLR, Germany 5Institute of Applied Physics, University of Bern, Switzerland



 Abstract

 TEIS (Terahertz and submm Limb Sounder) is a fooperation between European institutes, DLR, RAf and SRON, to build a three-channel balloon-borne therodyne spectrometer for atmospheric research. The optical front-end of the instrument consists of a spape 28x14 cm), calibration blackbody and the relay optics (common for the three channels). Beam spectrometer for atmospheric research. The optical front-end of the three channels is performed a polarizer. After splitting, the three beams enter the liquid helium optics (common for the three hoptical design and relay optics (common for the three bornes).

 In this poster we present the optical design and for poster we present the optical design and for TELIS. It is based on a phase-locked by is placed on the flat back surface of the elliptical is placed on the flat back surface of the elliptical for the three shaping and relaying of the beam is done bu.

 Big and validation of the optical design and for the shaping and relaying of the beam is done bu.

 Big and validation of the optics, as well as based on the flat back surface of the elliptical back surface of the date on the flat back surface of the elliptical back surface of the e



Cold optics design & amplitude-phase measurements

Layout of the SIR cold channel. In the test flight the SIR will operate in double sideband mode; SSB filter is replaced by a set of two plane mirrors (not shown in the picture).





Optical (dotted line) and quasioptical (solid line) trajectories (1/e field level @625GHz) in the cold channel. The curved mirrors (L1), (L3), (L4) and L5), the window (L2) and the integrated lensantenna (L6) are given schematically as thin vertical lines, representing thin lenses.



SIR cold channel phase distribution. Distance from the waist position – 110mm. Frequency - 600GHz.





Schematics of the 500-650 GHz channel optics. The telescope is rotated around the axis coinciding with the direction of the output beam. Wire grid polarizer and dichroic plate are used to separate this receiver from the two other frequency channels (not shown). The cold optics and mixer element are located inside the cryostat at the ambient temperature 4.2 K.



Calculated by GRASP far field 1-D vertical (elevation) cut and Azimuthally Collapsed Antenna Pattern (ACAP) at 625 GHz.

SIR channel optics design

Summary

> Integrated lens-antenna has been modeled using PILRAP. Far field amplitude beam measured at 600GHz is in a good agreement with
> The optics has been designed using 1-D MathCad tool and verified by GRASP. Amplitude-phase beam measurements of the cold optics
are performed at 600GHz. The measured beam waist is 2.25mm (within1% of the designed value), Gaussisity of the measured beam is
92.4%.
> Compact range (copy of the TELIS telescope with horizontally oriented slit source at the focus) is setup for the vertical beam profile
measurements. Preliminary tests indicate 0.4deg FWHM beam at 600GHz, in excellent agreement with the design. The corresponding
beamsize at tangent point is 1.6km.



## Acknowledgments

The authors thank M. van der Vorst for supplying PILRAP (Program for Integrated Lens and Reflector Antenna Parameters) and Jan Barkhof for assistance amplitude-phase in beam measurements.

For further information p.a.yagoubov@sron.rug.nl information please contact: